# cosmic microwave background physics

Heraeus summer school on cosmology, Heidelberg 2013



Centre for Astronomy Fak<mark>ultät für Physik und Ast</mark>ron<mark>omie, Universität Heide</mark>lberg

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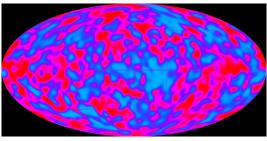






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#### cosmic microwave background



source: COBE observations

- the Universe is filled with radiation corresponding to a temperature of 2.726K
- small fluctuations of the temperature of the sky of order 10<sup>-5</sup>
- radiation from the formation of the first atoms

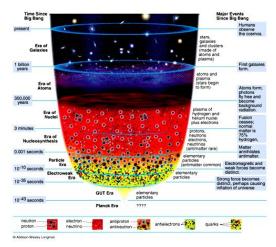
# thermal history of the universe

- temperature of fluids drop while universe expands
- 2 important stages

  - temperature is high enough to allow nuclear reactions  $\rightarrow$  big bang nucleosynthesis ( $z \simeq 10^{10}$ )

  - 2 temperature is high enough to ionise hydrogen  $\rightarrow$  cosmic microwave background ( $z \simeq 10^3$ )

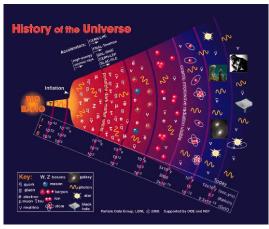
#### thermal history of the universe: overview



#### source: Addison-Wesley

СМВ

### thermal history of the universe: particle interactions



source: particle data group

СМВ

#### temperature and Hubble expansion

- Hubble expansion is an **adiabatic** process  $\delta Q = 0$
- adiabatic equation:  $V^{\kappa-1}T = \text{const}$  with adiabatic index  $\kappa \equiv c_p/c_V$
- early times: universe is filled with photons  $\kappa = 4/3$  (relativistic gas)

$$T \propto V^{-1/3} \propto a^{-1} \tag{1}$$

• late times: universe is filled with (dark) matter  $\kappa = 5/3$  (classical gas)

$$T \propto V^{-2/3} \propto a^{-2} \tag{2}$$

#### **Planck-spectrum for photons**

 photons in thermodynamic equilibrium are characterised by the Planck-spectrum

$$n(p,T) = \frac{g}{(2\pi\hbar)^3} \int_0^\infty \mathrm{d}p \frac{4\pi p^2}{\exp(\epsilon(p)/k_B T) - 1}$$
(3)

- Planck-spectrum depends only on temperature
- from the number density *n*(*p*, *T*) of photons we can compute number, energy and pressure by integration

$$n_{\gamma} = \frac{g_{\gamma}\zeta(3)}{\pi^2} \left(\frac{k_B T}{\hbar c}\right)^3, u_{\gamma} = \frac{g_{\gamma}\pi^2}{30} \frac{(k_B T)^4}{(\hbar c)^3}, p_{\gamma} = u_{\gamma}/3$$
(4)

- there are two polarisation states,  $g_{\gamma} = 2$
- pure magic:  $u_{\gamma} \propto a^{-4}$  (dilution and redshift), and at the same time:  $u \propto T^4$ , so  $T \propto a^{-1}$  as predicted from the adiabatic equation

# first atoms form

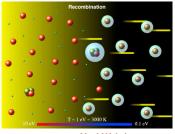


#### source: science kids

- at high temperatures, the reaction p + e<sup>−</sup> ↔ H + γ proceeds in both directions
- as the Universe expands, the temperature drops because of adiabatic cooling
- at low temperatures, the reaction only proceeds in the →-direction and atoms form

• this happens at  $\sim 10^4 {\rm K}$  roughly 300000 years after the big bang  $_{\rm Dorn \ Malte \ Schäfer}$ 

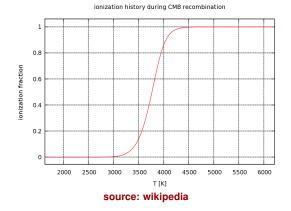
# photon propagation



source: Ned Wright

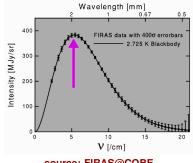
- while the Universe is hot, all atoms are ionised: photons scatter off electons and can't propagate
- Universe cools and atoms form: photons can travel freely and the Universe becomes transparent
- we see this radiation redshifted by 1000 today as the microwave background

#### formation of atoms



- fraction of neutral atoms is a steep function of temperature
- while the Universe cools down, the atoms form really fast

#### cosmic microwave background



source: FIRAS@COBE

- · atoms were produced in thermal equilibrium
- photons should follow a Planck-distribution
- redshifted by 1000 since then, from optical to microwave

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# **COsmic Background Explorer**

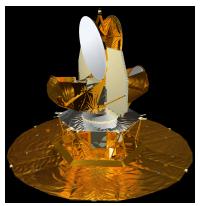


source: NASA



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#### Wilkinson Microwave Anisotropy Probe



source: NASA

WMAP-satellite

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### **Planck-surveyor**

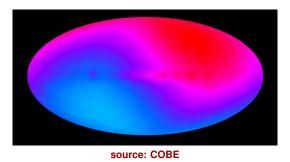


source: ESA

#### Planck-satellite

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#### **CMB** motion dipole



- the most important structure on the microwave sky is a dipole
- CMB dipole is interpreted as a relative motion of the earth
- CMB dispole has an amplitude of  $10^{-3}K$ , and the peculiar velocity is  $\beta = 371 \text{km/s/c}$

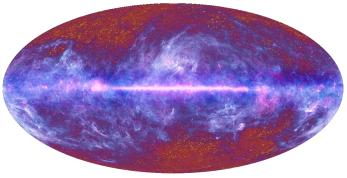
$$T(\theta) = T_0 \left(1 + \beta \cos \theta\right) \tag{5}$$

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# cosmic microwave background

- the temperature of the sky is not constant, but there are very small fluctuations
- the hot baryon plasma feels fluctuations in the distribution of (dark) matter by gravity
- at the point of (re)combination:
  - hydrogen atoms are formed
  - photons can propagate freely
- perturbations can be observed by two effects:
  - plasma was not at rest, but flowing towards a potential well → Doppler-shift in photon temperature, depending to direction of motion
  - plasma was residing in a potential well  $\rightarrow$  gravitational redshift

#### subtraction of motion dipole

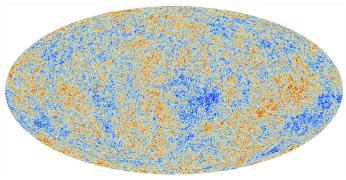


source: PLANCK

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CMB

# subtraction of Milky Way emission



source: PLANCK

#### what ...

...about those spots everywhere !?!

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# **CMB** angular spectrum

• analysis of fluctuations on a sphere: decomposition in  $Y_{\ell m}$ 

$$T(\theta) = \sum_{\ell} \sum_{m} T_{\ell m} Y_{\ell m}(\theta) \quad \leftrightarrow \quad T_{\ell m} = \int d\Omega \ T(\theta) Y_{\ell m}^*(\theta) \tag{6}$$

- spherical harmonics are an orthonormal basis system
- average fluctuation variance on a scale  $\ell \simeq \pi/\theta$

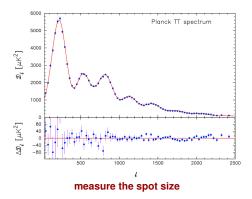
$$C(\ell) = \langle |T_{\ell m}|^2 \rangle \tag{7}$$

 averaging (...) is a hypothetical ensemble average. in reality, one computes an estimate of the variance,

$$C(\ell) \simeq \frac{1}{2\ell + 1} \sum_{m=-\ell}^{m=+\ell} |T_{\ell m}|^2$$
 (8)

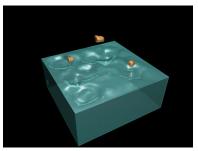


#### what about those spots?



- we compute the Fourier transformation and measure the angular size of the object (aka the wavelength)
- there's a peak in the spectrum at 2 degree: that's the size of the spots

#### sound waves in the plasma



superposition of sound waves

- processes in the early universe excite sound waves
- we see a superpositions of them in the cosmic microwave background
- there are temperature variations because the plasma is moving around in the sound wave

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#### standard ruler principle



distance estimate with a sniper scope

- estimate the distance to an object by measuring the angle under which it appears
- need to know the true physical size of the object

# standard ruler principle



#### trinity nuclear test, 16 milli-seconds after explosion

physical size: combine



- time since explosion
- 2 velocity of fireball
- distance: combine

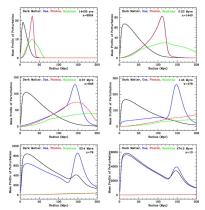


physical size

2 angular size



#### formation of baryon acoustic oscillations

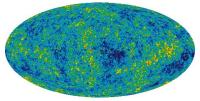


evolution of a single perturbation (source: Eisenstein, Seo and Hu (2005))

from a pointlike perturbation, a spherical wave travels in the photon-baryon-plasma

Björn Mate propagation stops when atoms form

### cosmic microwave background: standard ruler

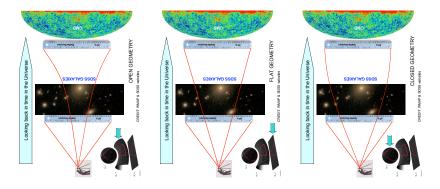


all-sky map of the cosmic microwave background, WMAP

- hot and cold patches of the CMB have a typical physical size, related to the horizon size at the time of formation of hydrogen atoms
- idea: physical size and apparent angle are related, redshift of decoupling known



#### standard ruler: measurement principle



- curvature can be well constrained
- assumption: galaxy bias understood, nonlinear structure formation not too important

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#### distance measures: comoving distance

- comoving distance  $\chi$  is the distance on a spatial hypersurface between the world lines of a source and the observer moving with the Hubble flow
- photon geodesics are defined by ds = 0 (Fermat's principle)
- therefore  $cdt = -ad\chi$  (from metric),  $d\chi = -cda/(a^2H)$

$$\chi = c \int_{a_e}^{a_a} \frac{\mathrm{d}a}{a^2 H(a)} \tag{9}$$

• complete analogy to conformal time  $d\eta = da/(a^2H)$ , such that  $\chi = c\eta$ 

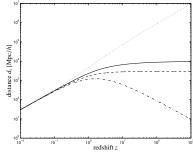
### distance measures: angular diameter distance

- angular diameter distance *d* is the distance infered from the angle under which a physical object appears
- physical cross section  $\Delta A$ , solid angle  $\Delta \Omega$ :

$$\frac{\Delta A}{4\pi a_{e\chi}^2} = \frac{\Delta\Omega}{4\pi}$$
(10)

• define d:  $d \equiv \sqrt{\frac{\Delta A}{\Delta \Omega}} = a_{e\chi}$ (11) CMB

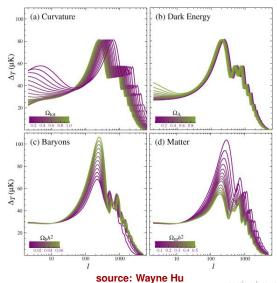
### relation between distance and redshift



cosmological distances vs. redshift z

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#### parameter sensitivity of the CMB spectrum



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# **CMB** simulator

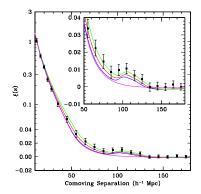
#### **CMB** simulator

http://www.strudel.org.uk/planck/

#### **Planck paper model**

http://planck.cf.ac.uk/news/make-your-own-planck-model

# baryon acoustic oscillations in the galaxies



pair density  $\xi(r)$  of galaxies as a function of separation r

 baryon acoustic oscillations: the (pair) density of galaxies is enhanced at a separation of about 100Mpc/h comoving

 idea: angle under which this scale is viewed depends on redshift Björn Matte Schäfer



### summary: microwave background

- · we can today observe the radiation from the formation of atoms
- the atoms formed at a temperature of 3000K at a redshift of 1000, and today



- temperature is 3K
- frequency is 160GHz
- 3 wave length is 3mm
- the optical light is shifted to microwaves by cosmological redshifting
- redshifting corresponds to adiabatic cool-down in the expansion



#### summary: spots in the CMB

- · the temperature has tiny fluctuations: there are spots in the CMB
- sound waves are excited in the plasma in the early Universe
- the sound waves travel until atoms form
- a standard ruler of size  $c_s \times \Delta t$  is established
- we observe this standard ruler under an angle of 1...2 degrees
- we know how far the cosmic microwave background is away, and have an integrated measure of the Hubble function